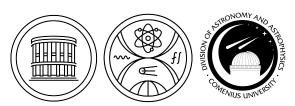
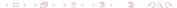
# A Tejútrendszer dióhéjban

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2021.08.09.



## **OUTLINE**

- Introduction
- ② GALACTIC MORPHOLOGY
- GALACTIC KINEMATICS
- GALACTIC DYNAMICS
- **(5)** Observations of the Milky Way
- 6 Farewell

### THE MILKY WAY GALAXY

### WHAT IS THE MILKY WAY?

A complex system of gravitationally bound stars, interstellar gas and dust and perhaps dark matter

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- The nomenclature of the Milky Way has its roots in the Greek mythology (just as the term galaxy)
- There are some interesting nomenclatures in our region, too:
  - The Path of Armies (Star-path of Prince Csaba)
  - The Path of the Fallen Souls
  - The Dewy Way (The Path of the Fairies)
  - The Fairy Silk Tissue
  - The Straw Way (The Way of the Gypsy, St. Peter's Way)



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- Shapley & Curtis (1921) The Great Debate until Hubble (1926) cooled them down
- From the 1920s observations of other galaxies



## OUR CURRENT KNOWLEDGE - A FACT SHEET

### THE RECIPE OF UNDERSTANDING OUR GALAXY

Observations of other galaxies + Statistical methods + Precise observational campaigns

| Туре           | Sbc-type barred spiral                | Gerhard (2002)                      |
|----------------|---------------------------------------|-------------------------------------|
| $R_{visible}$  | $\sim 31~{ m kpc}$                    | Lopez-Corredoira et al. (2018)      |
| $R_{DM}$       | $\sim 292~{ m kpc}$                   | Deason et al. (2020)                |
| R <sub>⊙</sub> | $\sim 8.122~{ m kpc}$                 | Gravity Collaboration et al. (2018) |
| Count of stars | 200-400 billion                       | many estimates                      |
| $M_{overall}$  | $1.18 \times 10^{12}~{\rm M}_{\odot}$ | Watkins et al. (2019)               |

### THE MORPHOLOGY OF THE GALAXY

Mapping and creating a model of a complex system of stars and interstellar material from the inside of such system is a challenging task.

#### IT CAN BE DONE VIA

- Observing galexies beyond the MW
- Creating abundance analyses
- Measuring parallaxes



## THE CURRENT MODEL - IN PICTURES

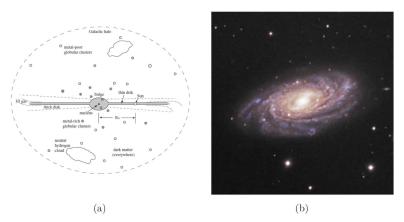


FIGURE: (a): Schema of the Milky Way. *Credit:* Sparke and Gallagher III (2007). (b): Galaxy NGC 3953. *Credit:* Wilson (2010).

Thick & thin disks

- •Not completely cylindrically symmetric, rather elliptic
- •Thin disk younger stars & interstellar medium
- $\bullet z_{\rm thin} \sim 350~(90)~{
  m pc}$
- $\bullet M_{\rm thin} \sim 6~(0.5) \times 10^{10}~{\rm M}_{\odot}$
- •Thick disk older stars
- $\bullet z_{\rm thick} \sim 1000~{
  m pc}$
- $\bullet M_{\rm thin} \sim 3 \times 10^{10} \ {\rm M}_{\odot}$



Thick & thin disks

- •Thin disk is composed of the metal-rich and young Population I stars with  $[Fe/H]\sim 0.02$
- $\bullet$  Thick disk contains metal-poor and older stars with  $[Fe/H] \sim 0.001$
- •Peculiar metallicity Population III discovered (Carollo et al., 2019)

Thick & thin disks

- •Spiral arms formed by O and B stars, H II region and stellar open clusters
- Quasi-stationary density wave in which a particular wave pattern stays in place, though the material forming the spiral does not reside in the location of the spiral
- •The spiral arms have a major role in star formation

- Thick & thin disks
- Interstellar gas & dust

- ullet most of ISM is located near the Galactic midplane, especially the molecular  $H_2$  and dust
- •Mass of HI:  $4 \times 10^9 \text{ M}_{\odot}$ Mass of H2:  $10^9 \text{ M}_{\odot}$
- •Gas density in the Solar neighbourhood is:  $0.04~{\rm M}_{\odot}{\rm pc}^{-3}$
- •Its main components are hydrogen (77%), molecules (17%) and ions (6%).

- Thick & thin disks
- Interstellar gas & dust
- Warp & flare

- •Almost every disk galaxy has a noticeable warp (infra HI observations)
- •Close approaches of galaxies or a collision with a dwarf galaxy could create a mechanical wave, which causes the warp of the Galactic disk
- •Flare increment of galactic disk scale height with the galactocentric distance

- Thick & thin disks
- Interstellar gas & dust
- Warp & flare
- Bulge & bar

- •In the central region of the Galactic disk the vertical scale height increases and creates the bulge.
- $\bullet M_{\text{bulge}} = 10^{10} \ M_{\odot}$
- •It contains stars of various metallicities
- •Bar is a substructure of the bulge
- •Its shape is caused by the interaction of the inflowing gas from the spiral arms and the radiation waves from the Galactic Center, which leads to changes in the orbits of stars in the bulge

- Thick & thin disks
- Interstellar gas & dust
- Warp & flare
- Bulge & bar
- Galactic center

- •Extinction of more than 30 magnitudes in visible wavelengths
- $\bullet \mathrm{Hot}$  and dense molecular clouds  $10^6~\mathrm{M}_\odot$
- •Stellar nucleus (0.2 pc)  $3 \times 10^7 \text{ M}_{\odot} \text{ pc}^{-3}$
- •Saggitarius A\*  $3.7 \times 10^6~{\rm M}_{\odot}$ ,  $16~{\rm R}_{\odot}$

- Thick & thin disks
- Interstellar gas & dust
- Warp & flare
- Bulge & bar
- Galactic center
- Stellar halo

- Contains globular clusters and field stars
- •Both are old metal-poor objects
- •Currently 150 known globular clusters with distances from 500 pc to 120 kpc from the GC (11 13 Gyr)
- •Stellar streams remnants of tidally shredded dwarf galaxies.

- Thick & thin disks
- Interstellar gas & dust
- Warp & flare
- Bulge & bar
- Galactic center
- Stellar halo
- Dark matter halo

- •Based on its gravitational influence over baryonic matter, there should be:
- $\bullet 5.4 \times 10^{11}~\text{M}_{\odot}$  of DM within 50~kpc from the GC
- $\bullet 1.9 \times 10^{12}~\text{M}_{\odot}$  of DM within 230~kpc from the GC
- •Which makes up about 95 % of the Galaxy

- Thick & thin disks
- Interstellar gas & dust
- Warp & flare
- Bulge & bar
- Galactic center
- Stellar halo
- Dark matter halo
- Satellite galaxies

- •About 60 known dwarf galaxies within 420 kpc radius
- •Most of them are old spheroidal ones
- •There are numerous traces of collisional and merger events with the Milky Way (the warp, stellar populations, new structures ...)

### DEFINING THE COORDINATE SYSTEMS

#### An appropriate system of coordinates

Takes advantage of the natural symmetry of the Galactic disk

Three coordinate systems are commonly used:

- Galactic coordinate system (l, b)
- Cylindrical coordinate system  $(R, \theta, z)$
- Cartesian coordinate system (x, y, z)

# THE GALACTIC COORDINATE SYSTEM

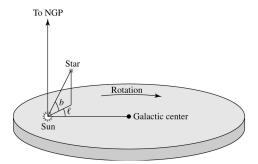


FIGURE: Credit: Carroll & Ostlie (2006)

### THE GALACTIC COORDINATE SYSTEM

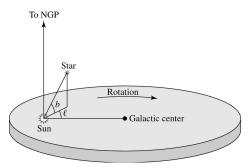
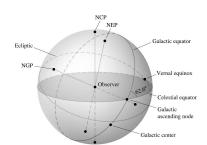
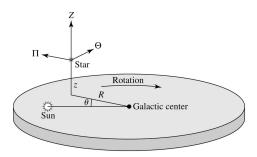


FIGURE: Credit: Carroll & Ostlie (2006)



## THE CYLINDRICAL COORDINATE SYSTEM



VELOCITY COMPONENTS

 $\Pi = \dot{R}, \quad \Theta = R\dot{\theta}, \quad Z = \dot{z}$ 

FIGURE: Credit: Carroll & Ostlie (2006)

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Is a point that is always centered on the Sun and moves on a perfectly circular orbit around the Galactic center (Or, the average motions of stars around the solar neighbourhood.)

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However, this values widely varies between authors

$$\Theta_{LSR} = 220 - 240 \text{ km s}^{-1}$$



### THE PECULIAR MOTIONS

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The relative velocity of a star to the LSR is called *peculiar velocity* 

- If we measure a star's velocity, in fact we measure the difference between the star's and Sun's peculiar velocity
- Overall peculiar velocity of the Sun (Solar motion) is

$$13.4 \; {\rm km \; s^{-1}}$$

towards the constellation Hercules



### THE MOTION OF STARS

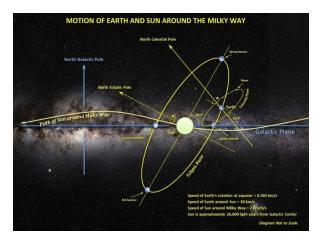


FIGURE: Credit: This blog



## THE MOTION OF STARS

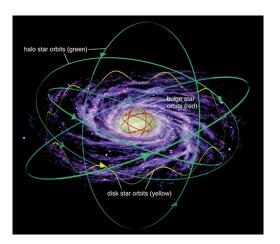


FIGURE: Credit: This lecture

### DIFFERENTIAL ROTATION

#### WHY?

The Galactic disk(s) does not rotate as a solid wheel, the orbital period varies with the Galactocentric distance



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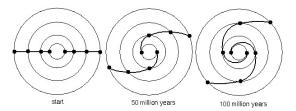


FIGURE: Differential rotation. Credit: This lecture

### DIFFERENTIAL ROTATION

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#### THE OORT'S CONSTANTS

This rotation could be described by the Oort's constants

$$A = (15.3 \pm 0.4) \text{ km s}^{-1} \text{ kpc}^{-1} ,$$
  
$$B = (-11.9 \pm 0.4) \text{ km s}^{-1} \text{ kpc}^{-1} .$$

where A describes the shearing motion, while B describes the vorticity, the Galactic rotation in the solar neighbourhood



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#### IF GALAXIES OBEY NEWTON'S LAW OF UNIVERSAL GRAVITATION

... there must be an excess of unseen mass



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#### UP TILL NOW

Dark matter remains unexplained



#### WHAT IF NEWTON WAS WRONG?

There would be no need for the dark matter. Maybe the gravity works in a different way ...

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- Klačka (2019) Proposal of the Generalized (and modified) gravitation based on the latest observations

### Introduction to Galactic Dynamics

 Replacing Newton's law with the Poisson's equation one can describe the relation between the gravitational force and mass as

$$\vec{\nabla}^2 \phi(\vec{r}) = 4\pi G \rho(r)$$

• it is assumed that the Poisson's equation holds on the scale of galaxies and the Universe

#### IN ORDER TO STUDY DYNAMICS OF GALAXIES

the mass density distribution  $\rho(r)$  and the gravitational potential  $\phi(r)$  must be determined



### WHY ARE DENSITY AND POTENTIAL SUCH CRUCIAL?

ullet A discrete and collisionless system of N point masses could be defined by a six-dimensional phase-space density (distribution function) :

$$f(\vec{r},\dot{\vec{r}},t) \ \mathrm{d}^3 \vec{r} \ \mathrm{d}^3 \dot{\vec{r}}$$

• Collisionless system - time derivative must be zero as the flow in the phase-space is incomprehensible:

$$\dot{f}(\vec{r},\dot{\vec{r}},t) = 0$$

#### THE VLASOV EQUATION

completely specifies the evolution of a collisionless system



### WHY ARE DENSITY AND POTENTIAL SUCH CRUCIAL?

• Once the distribution function is obtained, the density could be obtained straightforward by integrating over the velocity space:

$$\rho(\vec{r},t) = \iiint f(\vec{r},\dot{\vec{r}},t) \, \mathrm{d}^3\dot{\vec{r}}$$

while the total mass of the system is determined from:

$$M(t) = \iiint f(\vec{r},t) \ \mathrm{d}^3\vec{r} = \int \mathrm{d}^3\vec{r} \int \mathrm{d}^3\dot{\vec{r}} \ f(\vec{r},\dot{\vec{r}},t)$$

#### The gravitational potential

must be constructed reasonably



### EXAMPLE - SIMPLE SPHERICAL POTENTIAL

To present how to obtain the gravitational force, we will derive Newton's universal gravitation. Considering a simple spherical density distribution  $\rho(r)$  we obtain the Poisson's equation as follows:

$$\frac{1}{r^2} \frac{\partial}{\partial r} \left( r^2 \frac{\partial \phi}{\partial r} \right) = 4\pi G \rho(r) ,$$

from which we obtain

$$r^2 \frac{\partial \phi}{\partial r} = 4\pi G \int_0^r \rho(r) r^2 \, dr = GM(r),$$

where M(r) is the total enclosed mass. Recalling that acceleration is the negative gradient of the potential we obtain Newton's gravitational force:

$$\vec{F}(\vec{r}) = -\vec{\nabla}\phi(\vec{r}) = -\frac{\mathrm{d}\phi}{\mathrm{d}r}r = -\frac{GM(r)}{r^2}r \ , \label{eq:free}$$

where  $\vec{r}$  is position vector with respect to the center of the mass and r is the unit position vector. While the gravitational potential which obeys both Newton's law and the Poisson's equation is

$$\phi(\vec{r}) = -\frac{GM(r)}{r} \ .$$

### THE GALACTIC ROTATION CURVE

 The circular velocity is determined from the relation

$$v_{\rm c}(R) = \sqrt{R \frac{\partial}{\partial R} \left(\phi_{\rm Total}\right)}$$

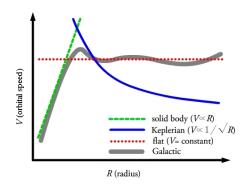
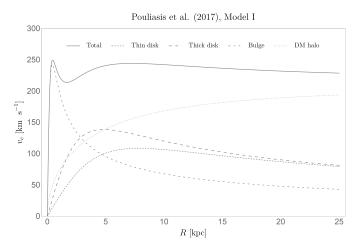


FIGURE: Rotation curves. Credit: This website.

## How does a real $\phi_{\text{total}}$ look like?

$$\begin{split} \phi_{\rm Thick\; disk}(R,z) &= \frac{-GM_{\rm Thick\; disk}}{\left[R^2 + \left(a_{\rm Thick\; disk} + \sqrt{z^2 + b_{\rm Thick\; disk}^2}\right)^2\right]^{1/2}}\;,\\ \phi_{\rm Thin\; disk}(R,z) &= \frac{-GM_{\rm Thin\; disk}}{\left[R^2 + \left(a_{\rm Thin\; disk} + \sqrt{z^2 + b_{\rm Thin\; disk}^2}\right)^2\right]^{1/2}}\;,\\ \phi_{\rm Bulge}(R) &= \frac{-GM_{\rm Bulge}}{\left(R^2 + b_{\rm Bulge}^2\right)^{1/2}}\;,\\ \phi_{\rm DM\; halo}(R) &= \frac{-GM_{\rm DM\; halo}}{R} - \frac{M_{\rm DM\; halo}}{1.02a_{\rm DM\; halo}} \times \left[\frac{-1.02}{1 + (R/a_{\rm DM\; halo})^{1.02}} + \ln\left(1 + \left(\frac{R}{a_{\rm DM\; halo}}\right)^{1.02}\right)\right]_{\rm DM\; halo}^{100}\;. \end{split}$$

## How does a real $\phi_{\text{total}}$ look like?



 $\overline{F}IGURE$ : Rotation curves produced from potential model I of Pouliasis et al. (2017).

### MOND AND MOG

# MODIFIED NEWTONIAN DYNAMICS

the force due to the Newtonian gravity for very small values of acceleration  $|\vec{a}|$  is scaled as

$$m\frac{|\vec{a}|}{a_0}\vec{a} = g_{bar}$$

(Milgrom, 1983)

#### MODIFIED GRAVITATION

keeps Newton's second law in it's original form, but changes the gravitational force

$$m\vec{a} = \frac{g_{bar}}{1 - e^{-\sqrt{g_{bar}/g_{\dagger}}}}$$

(McGaugh et al., 2016), (Klačka, 2019)

$$g_{bar} = - \frac{GMm}{|\vec{r}|^3} \vec{r} \; ; \qquad a_0 = g_\dagger = (1.2 \pm 0.02) \times 10^{-10} \; \mathrm{m \; s^{-2}}$$

### OBSERVATIONAL SURVEYS

### Anders et al. (2019)

"Galactic astrophysics is currently in a similar phase as geography was in the 15th century: large parts of the Earth were unknown to contemporary scientists, only crude maps of most of the known parts of the Earth existed, and even the orbit of our planet was still under debate. Nowadays, major parts of the Milky Way are still hidden by thick layers of dust, but we are beginning to discover and to map our Galaxy in a much more accurate fashion by virtue of dedicated large photometric, astrometric, and spectroscopic surveys."

Hipparcos

- •The first astrometric telescope in space
- •Astrometric and photometric data from 1989 to 1993
- •Overall 118 218 stars observed
- $\bullet$ Tycho 2 Catalogue (Hog et al., 2000) 2.5 million stars 99 % of stars down to magnitude 11
- •The widest catalogue of stars at the end of the millennium.

- Hipparcos
- 2MASS

- •Ground-based observational campaign
- •Two dedicated telescopes covered 99% of the celestial sphere in near-infrared bandpass
- •4.1 million FITS images of the entire night sky and over 461 million objects with astrometry and photometry, including stars, brown dwarfs, minor objects in the Solar System, gas nebulae and stellar clusters (Skrutskie et al., 2006)

- Hipparcos
- 2MASS
- WISE

- Successor of 2MASS
- •Its goals were to search for the origin of planets, stars and galaxies and to create an infrared map of the Universe
- •Its catalogue contains (i) positions, apparent motions and flux variability statistics for more than 747 million objects, (ii) a photometry database containing 42 billion time-tagged measurements and (iii) an image atlas with  $\sim 18000$  FITS images.

- Hipparcos
- 2MASS
- WISE
- SDSS

- •Its goal is to create a map of the Universe and the distribution of the luminous and non-luminous matter.
- •Still operating, from 2000.
- •So far imaged 8000 square degrees of the sky, measured spectra of more than 700 000 objects and mapped one million galaxies and 100 000 quasars
- •Currently is observing with extended cosmological precision, it creates the first spatially resolved maps of galaxies and its expanding the spectroscopic survey of the Milky Way galaxy.

- Hipparcos
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#### ACHIEVEMENTS

•All together the SDSS is **probably the most successful survey** in the history of astronomy.

- Hipparcos
- 2MASS
- WISE
- SDSS
- APOGEE

- •Is a part of the SDSS survey
- •Infrared spectroscopy observations of Galactic disk and bulge
- $\bullet \mbox{With}~163000$  targets, it is the largest scale spectroscopic survey of our Galaxy
- •The upcoming Data release will include more than 430000 stars.

### THE BILLION STAR SURVEYOR

#### GAIA

The Global Astrometric Interferometer for Astrophysics was launched in 2013 as a successor of Hipparcos

- Instruments: two identical optical telescopes, a radial velocity spectrometer, blue and red photometers.
- The payload features a focal plane array (0.5 m x 1 m) with 106 CCD detectors containing nearly 1 billion pixels, making it the largest digital camera ever used in space.
- Was planned for a five-year mission. However, an extension was approved until the end of 2022.

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#### DEFINITE CATALOGUE OF THE FORESEEABLE FUTURE

Data collected by Gaia throughout its mission will be used to eventually build the most accurate three-dimensional map of the positions, motions, and chemical composition of stars in our Galaxy

### Data Release 3

EDR3 ( $3^{RD}$  of December 2020)

The full DR3 in will released in the  $2^{\mbox{nd}}$  half of 2021

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The full DR3 in will released in the  $2^{nd}$  half of 2021

- ullet positions, parallaxes and proper motions for  $\sim 1.5$  bil. stars
- photometry for 1.8 billion sources
- about 1.5 mil. celestial reference frame sources (quasars)
- cross-matches between EDR3 and: DR2, Hipparcos-2, Tycho-2, 2MASS PSC, SDSS DR13, Pan-STARRS1 DR1, SkyMapper DR1, GSC 2.3, APASS DR9, RAVE DR5, allWISE, and URAT-1 data on the other hand

### THE ULTIMATE GAIA CATALOGUE

Gaia catalogue (end of the 2020's)

The release date is still unknown as additional mission extensions may occur

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- all available variable, binary and multiple stars solutions
- additional astrophysical parameters for stars, unresolved binaries, galaxies and quasars
- exoplanet catalogue
- all epoch and transit data for all sources

# Thank you for your attention

"But high above there is the cloudless sky of purity and light and grandeur too, as frail yet firm as everything that's true."

- Dezső Kosztolányi, Dawnstruck

